

# The Dual Cascade in Bounded and Unbounded Two-Dimensional Fluids

John C. Bowman and Chuong V. Tran

<http://www.math.ualberta.ca/~bowman/talks>

Department of Mathematical and Statistical Sciences, University of Alberta,  
Edmonton, Alberta, Canada

# Outline

## 1 2D Turbulence

- A KLB Theory
- B Balance Equation
- C Bounded 2D Turbulence

## 2 Numerical Simulations

- A Subgrid Models
- B Hyperviscosity
- C Supergrid Models

## 3 Conclusions

## 2D Turbulence

- Navier–Stokes equation for **vorticity**  $\omega = \hat{z} \cdot \nabla \times \mathbf{u}$ :

$$\frac{\partial \omega}{\partial t} + \mathbf{u} \cdot \nabla \omega = -\mathcal{D}\omega + f,$$

where  $\mathcal{D} = -\nu \nabla^2$  represents molecular dissipation.

- In Fourier space:

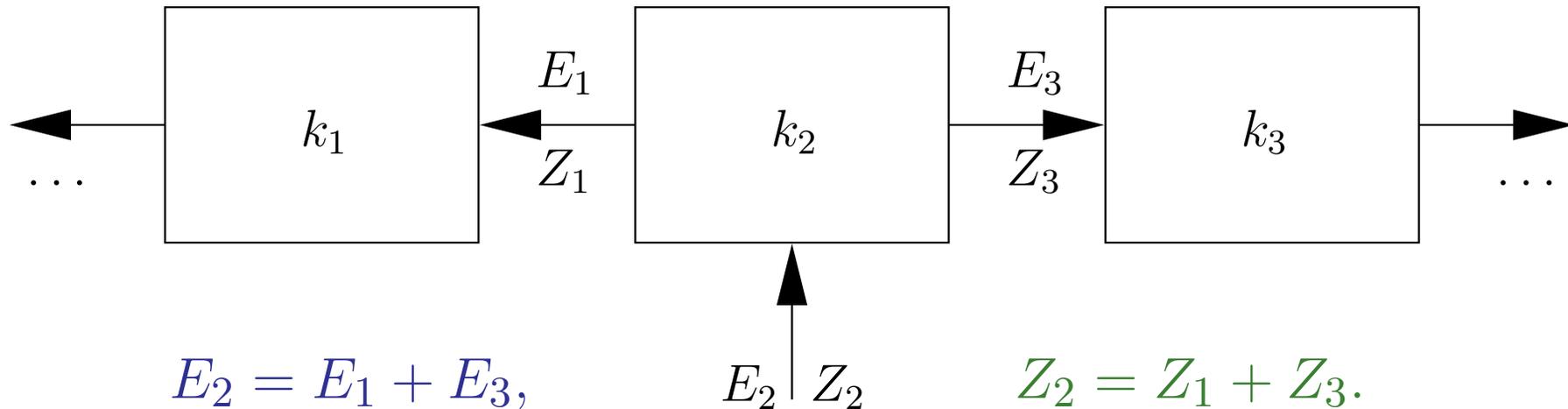
$$\frac{\partial \omega_{\mathbf{k}}}{\partial t} = S_{\mathbf{k}} - D_{\mathbf{k}} \omega_{\mathbf{k}} + f_{\mathbf{k}},$$

where  $D_{\mathbf{k}} = \nu k^2$ .

- The steady-state energy spectrum  $E(k) = \frac{1}{2} \sum_{|\mathbf{k}|=k} \frac{|\omega_{\mathbf{k}}|^2}{k^2}$  is of physical interest.

# KLB Theory

- Energy  $E = \frac{1}{2} \sum_k \frac{|\omega_k|^2}{k^2}$  and enstrophy  $Z = \frac{1}{2} \sum_k |\omega_k|^2$  are conserved.



- [Fjørtoft 1953]: energy cascades to large scales and enstrophy cascades to small scales.
- [Kraichnan 1967], [Leith 1968], and [Batchelor 1969] (KLB):  
 $k^{-5/3}$  inverse energy cascade on large scales,  
 $k^{-3}$  direct enstrophy cascade on small scales.

- Let  $s^2 = \overline{\sum_k f_k \omega_k^*} / \overline{\sum_k f_k \frac{\omega_k^*}{k^2}}$  be the ratio of mean **enstrophy** to **energy** injection.
- Typically,  $s$  will lie within the band of forced wavenumbers.
- Multiply the energy equation

$$\frac{1}{2k^2} \frac{\partial |\omega_k|^2}{\partial t} + D_k \frac{|\omega_k|^2}{k^2} = S_k \frac{\omega_k^*}{k^2} + f_k \frac{\omega_k^*}{k^2}$$

by  $s^2$  and subtract the enstrophy equation

$$\frac{1}{2} \frac{\partial |\omega_k|^2}{\partial t} + D_k |\omega_k|^2 = S_k \omega_k^* + f_k \omega_k^*$$

$\Rightarrow$  steady-state **balance equation** [Tran & Bowman 2003]:

$$\sum_{k=1}^s (s^2 - k^2) D_k E(k) = \sum_{k=s}^{\infty} (k^2 - s^2) D_k E(k).$$

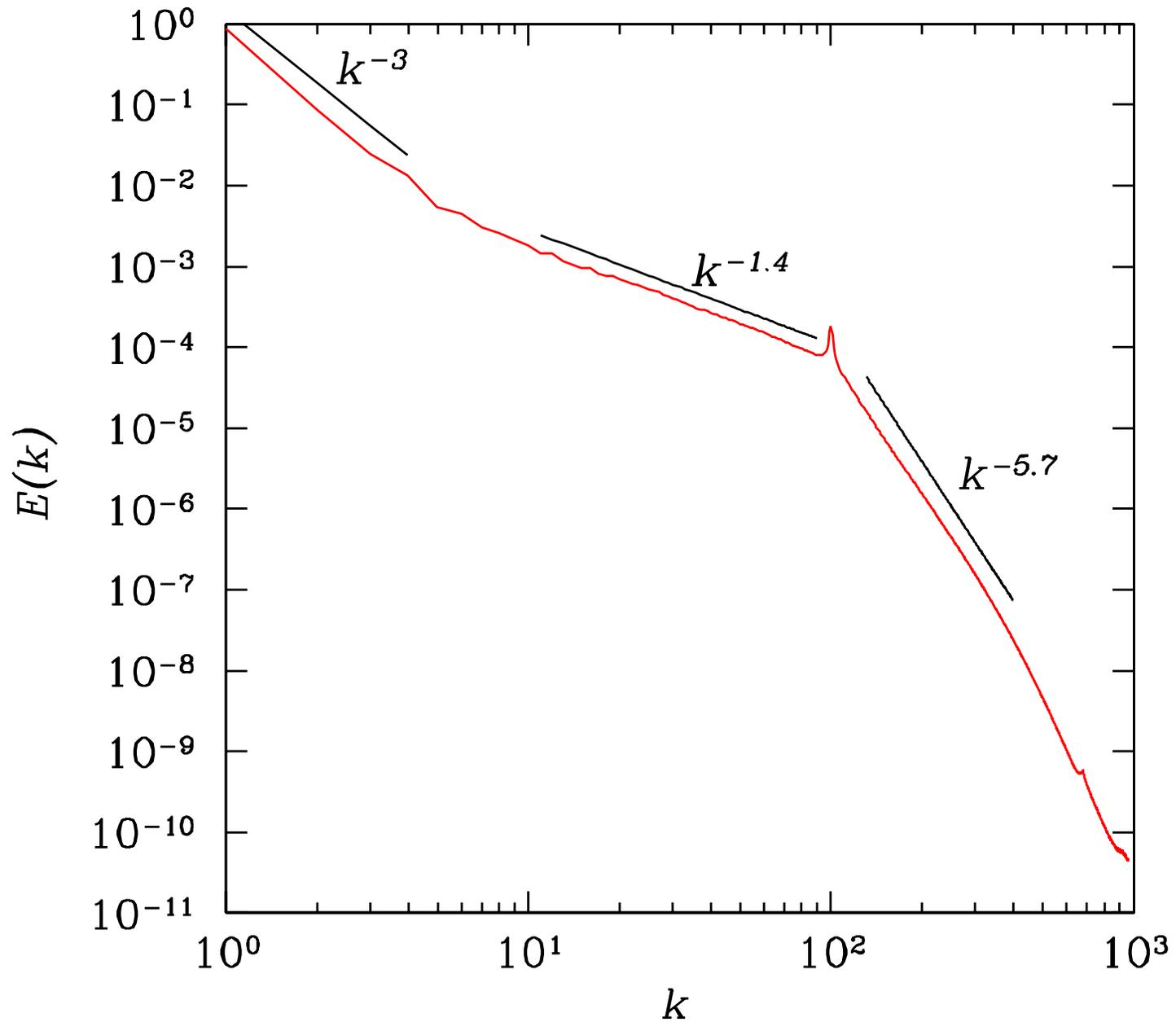
# Balance Equation

- Small and large scale dynamics are intricately coupled:

$$\sum_{k=1}^s (s^2 - k^2) D_k E(k) = \sum_{k=s}^{\infty} (k^2 - s^2) D_k E(k).$$

- Explains the discrepancy between the enstrophy-range KLB prediction  $E(k) \sim k^{-3}$  and the steep  $\sim k^{-5}$  spectrum typically seen in numerical simulations.
- **Unbounded domain:** everlasting inverse energy cascade.
- **Bounded domain:** upscale energy cascade is halted at the lowest wavenumber.
- Lower spectral boundary acts in effect as an external forcing.

# Large-scale direct cascade?



- Energetic reflections at the lower spectral boundary eventually lead to a large-scale **direct** “cascade.”
- This would agree with the large-scale  $k^{-3}$  spectra seen numerically [Borue 1994] and observed in the atmosphere [Lilly & Peterson 1983].
- [Tran & Bowman 2003]: In a bounded domain, the two inertial range exponents **must sum to  $-8$**  (at high Reynolds number).
- Large-scale  $k^{-3}$  spectrum  $\Rightarrow$  a small-scale  $k^{-5}$  spectrum.
- Consistent with rigorous [Tran & Shepherd 2002] constraint: the spectrum must be **at least as steep as  $k^{-5}$** .

## Bounded 2D Turbulence

- Q. How do the energy balances associated with the hypothetical steady-state energy spectrum

$$E(k) = A \begin{cases} k^{-\alpha} & \text{if } k_0 \leq k < s, \\ s^{\beta-\alpha} k^{-\beta} & \text{if } s \leq k \leq k_T \end{cases}$$

behave in the limit  $k_0 \rightarrow 0^+$ ,  $k_T \rightarrow \infty$ ?

- The energy dissipation would be equal to

$$\epsilon = 2\nu A s^{3-\alpha} \left( \frac{1}{3-\alpha} + \frac{1}{\beta-3} \right) \quad (\alpha < 3, \beta > 5).$$

- Apply steady-state constraint  $\alpha + \beta = 8$   
[Tran & Bowman 2003].

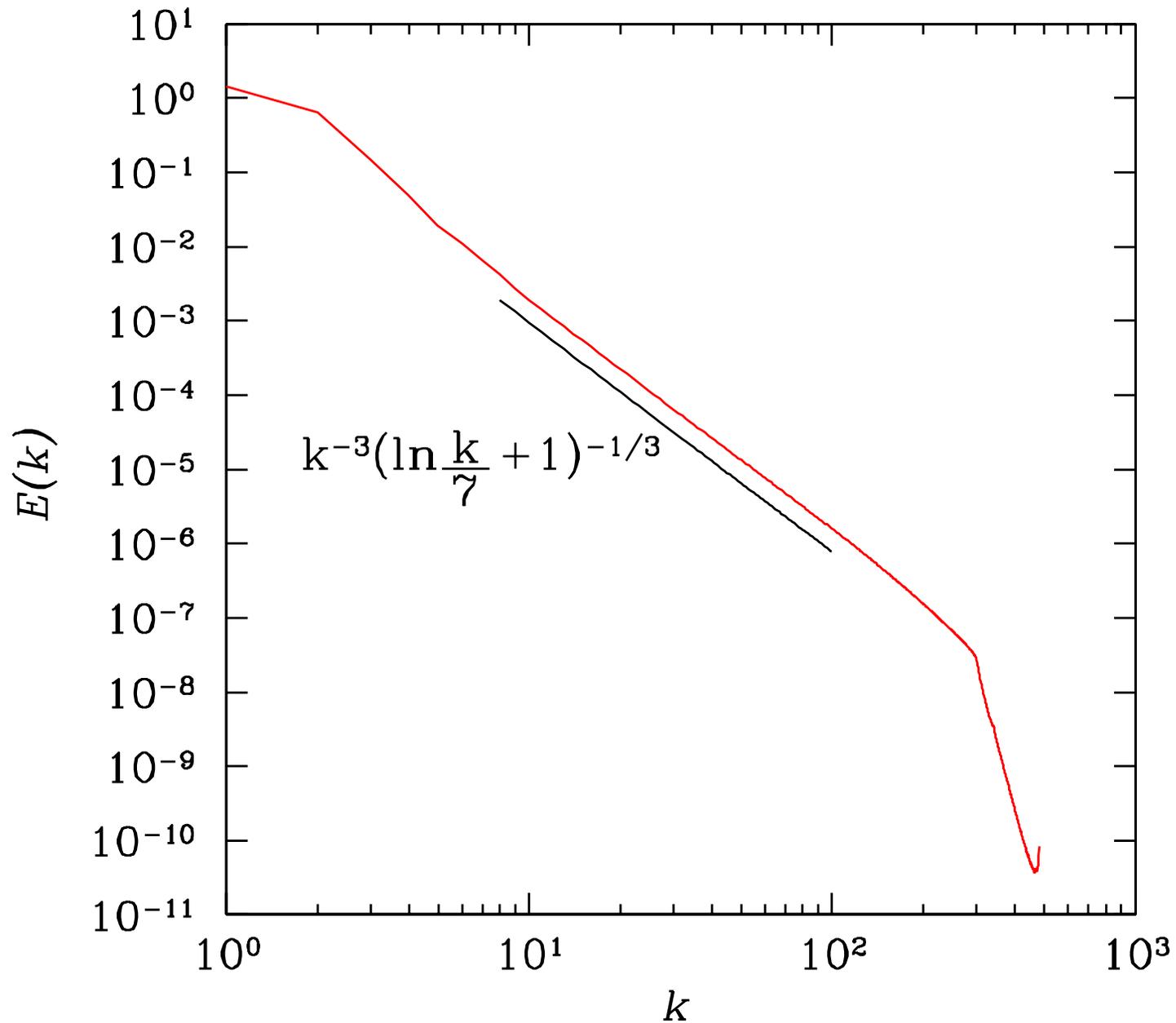
- Let  $\delta = 3 - \alpha = \beta - 5$  :

$$\epsilon = 2\nu A s^\delta \left( \frac{1}{\delta} + \frac{1}{2 + \delta} \right).$$

- If  $\lim_{\nu \rightarrow 0^+} A$  is finite then  $\lim_{\nu \rightarrow 0^+} \delta = 0$ .

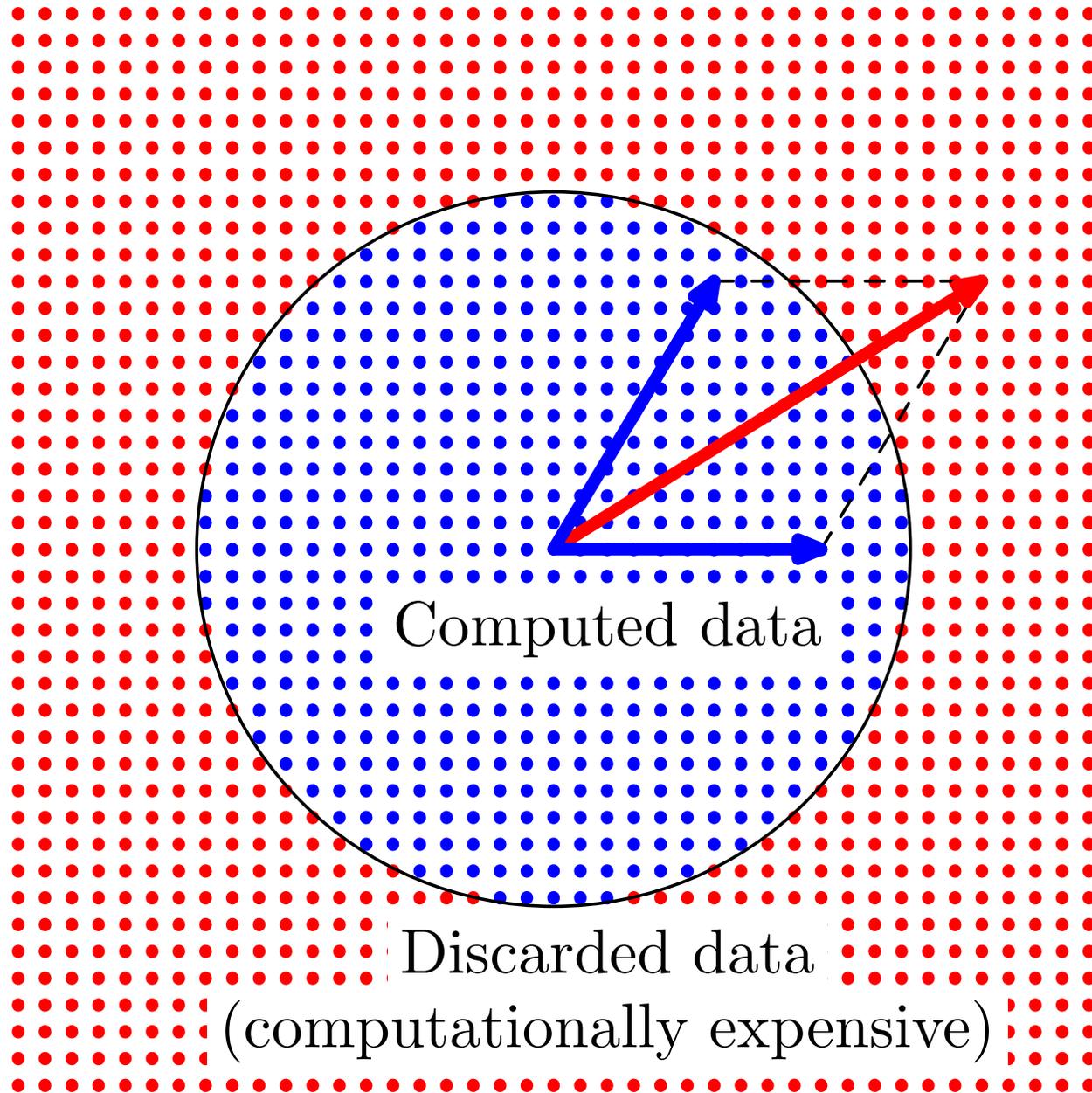
- That is,  $\lim_{\nu \rightarrow 0^+} \alpha = 3$  and  $\lim_{\nu \rightarrow 0^+} \beta = 5$ .
- **Claim:** steady-state high-resolution bounded numerical simulations, forced at an intermediate wavenumber, approach this limit.
- However, this says nothing about the **quasi-steady state** in an **unbounded domain** discussed by KLB (open problem).

# Direct $k^{-3}$ enstrophy Cascade



Zero dissipation for  $3 < k < 300$ .

# Subgrid Models



Must model the effect of red region on blue region.

# Hyperviscous Subgrid Model

- It is customary to replace interactions with missing small-scale modes by a **hyperviscous** term:

$$\frac{\partial \omega_{\mathbf{k}}}{\partial t} = \tilde{S}_{\mathbf{k}} - \nu k^2 \omega_{\mathbf{k}} - \nu_h k^n \omega_{\mathbf{k}} + f_{\mathbf{k}};$$

$\tilde{S}_{\mathbf{k}}$  accounts for the interactions involving only retained modes.

- It is often argued that this modification does not affect the large-scale dynamics.
- Inverse cascade  $\Rightarrow$  propagation of small-scale (mis)information **back to the large scales?**

# Supergrid Models

- A **hypoviscous** large-scale damping is typically added to thwart the upscale energy cascade:  $D_k = \nu_0 k^{-p} + \nu k^2$ , where  $p > 0$ .
- However, this **contaminates** the desired  $k^{-5/3}$  energy-range spectrum.
- An artificial large-scale damping can cause spurious energetic reflections or **bottleneck effects**.
- We need a large-scale **supergrid** model that takes energy out of the large scales in a realistic way (analogous to a small-scale **subgrid** model).
- Such a model should be compatible with the Kolmogorov's Ansatz of self-similar (wavenumber-independent) energy transfer.

# Enstrophy Transfer

- Recall

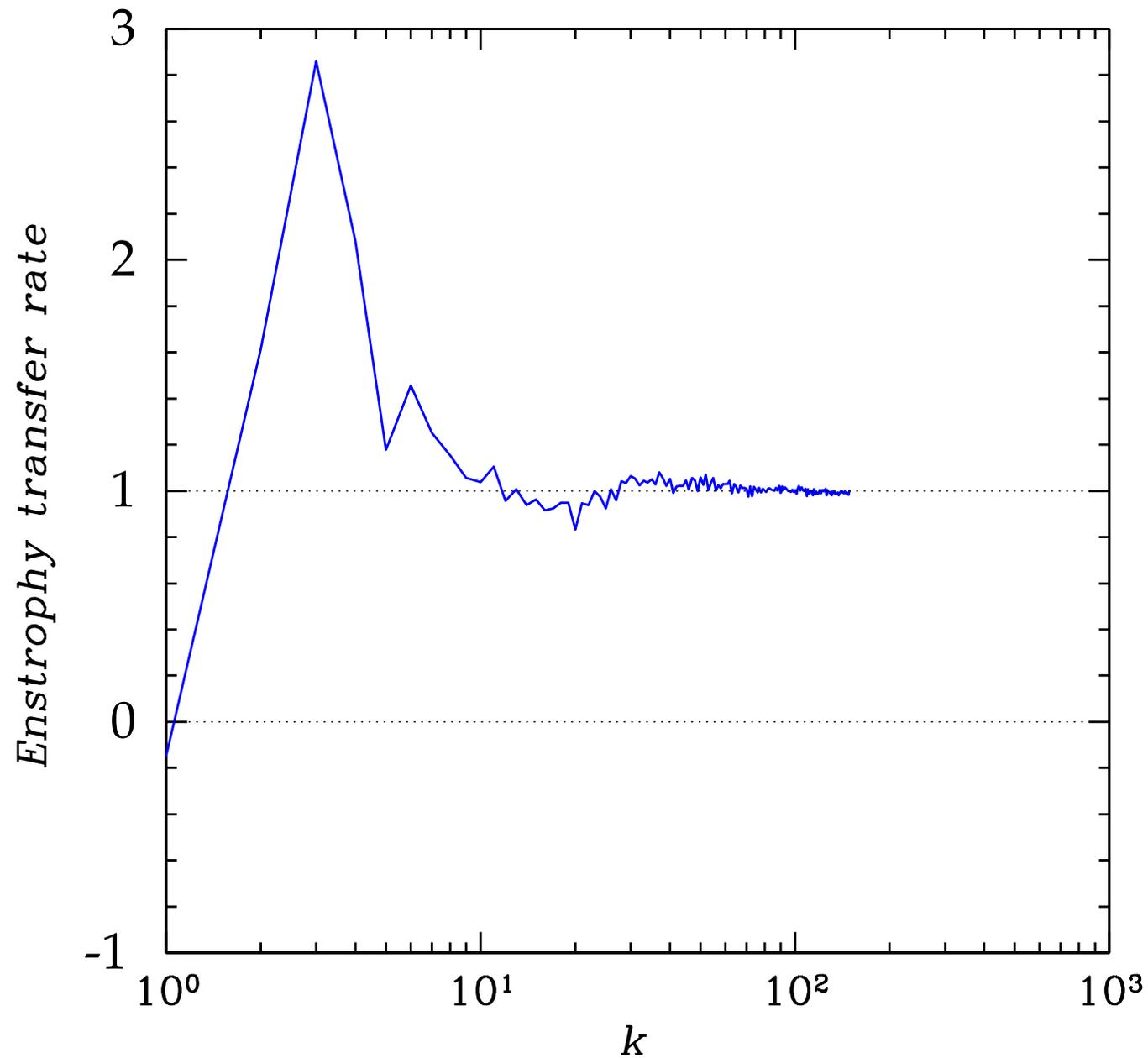
$$\frac{\partial \omega_{\mathbf{k}}}{\partial t} = \tilde{S}_{\mathbf{k}} - \nu k^2 \omega_{\mathbf{k}} + f_{\mathbf{k}}.$$

- Write  $\tilde{S}_{\mathbf{k}} = \sum_{\mathbf{p}} M_{\mathbf{k},\mathbf{p}} \omega_{\mathbf{p}} \omega_{\mathbf{k}-\mathbf{p}}$  in terms of the nonlinear mode-coupling coefficient:  $M_{\mathbf{k},\mathbf{p}} = \frac{\hat{\mathbf{z}} \cdot \mathbf{k} \times \mathbf{p}}{p^2}$ .

- The forward enstrophy transfer  $F_k$  through a wavenumber  $k$  in 2D can be computed as a **restricted convolution**:

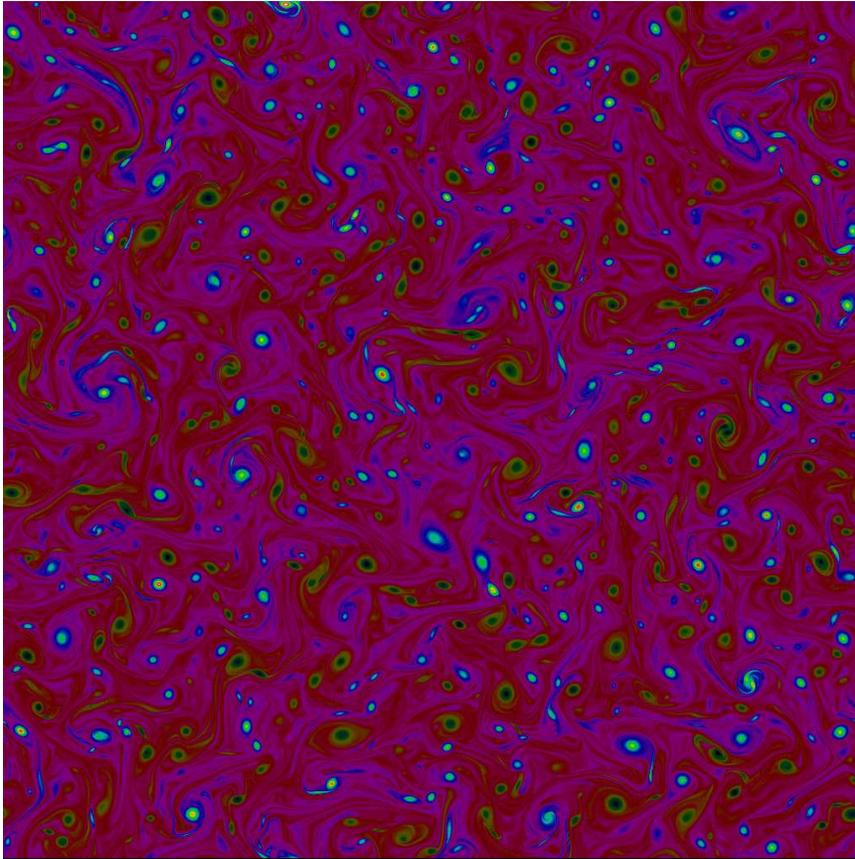
$$F_k = \operatorname{Re} \sum_{\substack{|\mathbf{k}|=k \\ |\mathbf{p}|<k \\ |\mathbf{k}-\mathbf{p}|<k}} M_{\mathbf{k},\mathbf{p}} \omega_{\mathbf{p}} \omega_{\mathbf{k}-\mathbf{p}} \omega_{\mathbf{k}}^* + 2 \operatorname{Re} \sum_{\substack{|\mathbf{k}|=k \\ |\mathbf{p}|<k \\ |\mathbf{k}-\mathbf{p}|>k}} M_{\mathbf{p},\mathbf{k}-\mathbf{p}} \omega_{\mathbf{p}} \omega_{\mathbf{k}-\mathbf{p}} \omega_{\mathbf{k}}^*.$$

# Self-Similarity of Enstrophy Transfer

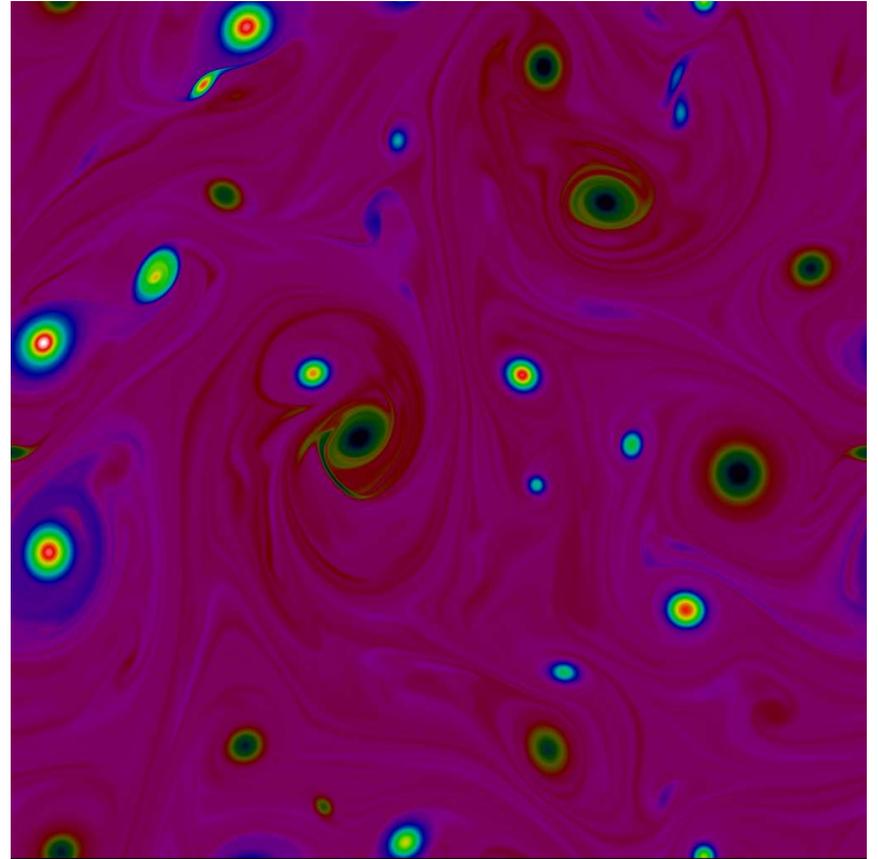


Enstrophy transfer rate  $F_k$  for  $k < 150$ .

# 2D Decaying Turbulence



early stage



late stage

## Conclusions

- A **direct** large-scale  $k^{-3}$  “cascade” resulting from reflections at the lower spectral boundary provides a physical explanation for numerically observed small-scale  $k^{-5}$  spectra.
- **Subgrid** and **supergrid** models could be used together to attempt to verify the KLB theory for unbounded fluids.
- We propose a **Self-Similar Turbulent Subgrid Model** based on Kolmogorov’s idea of scale-independent transfer.
- This may require the development of a **fast restricted convolution**.
- A proper subgrid model should account for both turbulent **damping** and **backscatter** effects.
- Decaying turbulence may involve a **spatial self-similarity**.

# References

- [Batchelor 1969] G. K. Batchelor, *Phys. Fluids*, **12** II:233, 1969.
- [Borue 1994] V. Borue, *Phys. Rev. Lett.*, **72**:1475, 1994.
- [Fjørtoft 1953] R. Fjørtoft, *Tellus*, **5**:225, 1953.
- [Kraichnan 1967] R. H. Kraichnan, *Phys. Fluids*, **10**:1417, 1967.
- [Leith 1968] C. E. Leith, *Phys. Fluids*, **11**:671, 1968.
- [Lilly & Peterson 1983] D. K. Lilly & E. L. Peterson, *Tellus*, **35A**:379, 1983.
- [Tran & Bowman 2003] C. V. Tran & J. C. Bowman, *Physica D*, **176**:242, 2003.
- [Tran & Shepherd 2002] C. V. Tran & T. G. Shepherd, *Physica D*, **165**:199, 2002.