

## WAVE EQUATIONS – EXPLICIT FORMULAS

In this lecture we derive the representation formulas for the wave equation in the whole space:

$$\square u \equiv u_{tt} - \Delta u = 0, \quad \mathbb{R}^n \times (0, \infty); \quad u(x, 0) = g(x), \quad u_t(x, 0) = h(x). \quad (1)$$

It turns out that the properties of the solutions depend on the dimension. More specifically, there are three cases:  $n = 1$ ,  $n > 1$  even;  $n > 1$  odd. We will discuss in detail the three representative cases:  $n = 1, 2, 3$  (the order is actually  $n = 1, 3, 2$ , for reasons that will be clear soon).

1.  $n = 1$ .

We consider the 1D wave equation

$$u_{tt} - u_{xx} = 0, \quad \mathbb{R} \times (0, \infty); \quad u(x, 0) = g(x), \quad u_t(x, 0) = h(x). \quad (2)$$

This equation can be solved via the following change of variables:

$$\xi = x + t; \quad \eta = x - t, \quad (3)$$

and to make things clearer we set  $\tilde{u}(\xi, \eta) = u(x, t)$ .

With this change of variable we compute

$$u_t = \tilde{u}_\xi \xi_t + \tilde{u}_\eta \eta_t = \tilde{u}_\xi - \tilde{u}_\eta \quad (4)$$

$$u_{tt} = \tilde{u}_{\xi\xi} - 2\tilde{u}_{\xi\eta} + \tilde{u}_{\eta\eta} \quad (5)$$

$$u_x = \tilde{u}_\xi + \tilde{u}_\eta \quad (6)$$

$$u_{xx} = \tilde{u}_{\xi\xi} + 2\tilde{u}_{\xi\eta} + \tilde{u}_{\eta\eta} \quad (7)$$

Therefore

$$u_{tt} - u_{xx} = 0 \iff \tilde{u}_{\xi\eta} = 0 \iff \tilde{u}(\xi, \eta) = \phi(\xi) + \psi(\eta) \iff u(x, t) = \phi(x + t) + \psi(x - t). \quad (8)$$

Now using the initial values we have

$$\phi(x) + \psi(x) = g(x); \quad \phi'(x) - \psi'(x) = h(x) \quad (9)$$

which yields

$$u(x, t) = \frac{1}{2} [g(x + t) + g(x - t)] + \frac{1}{2} \int_{x-t}^{x+t} h(y) \, dy. \quad (10)$$

This is d'Alembert's formula.

**Theorem 1.** Assume  $g \in C^2(\mathbb{R})$ ,  $h \in C^1(\mathbb{R})$ , define  $u$  by

$$u(x, t) = \frac{1}{2} [g(x + t) + g(x - t)] + \frac{1}{2} \int_{x-t}^{x+t} h(y) \, dy. \quad (11)$$

Then

i.  $u \in C^2(\mathbb{R} \times [0, \infty))$ ;

ii.  $u_{tt} - u_{xx} = 0$  in  $\mathbb{R} \times (0, \infty)$ ;

iii.  $u$  takes the correct boundary values:

$$\lim_{\substack{(x,t) \rightarrow (x_0,0) \\ t > 0}} u(x, t) = g(x_0); \quad (12)$$

$$\lim_{\substack{(x,t) \rightarrow (x_0,0) \\ t > 0}} u_t(x, t) = h(x_0). \quad (13)$$

**Proof.** The proof is by direct calculation and is left as an exercise. □

**Remark 2.** It is easy to generalize the above theorem to the case

$$g \in C^k, \quad h \in C^{k-1} \implies u \in C^k. \quad (14)$$

But in general  $u$  cannot be smoother. For example, consider the case  $h = g'$ , then  $u(x, t) = g(x + t)$ . It is clear that  $u$  cannot have better regularity than  $g$ .

**Remark 3.** As we have shown in the exercises after the lecture on distributions, the formula

$$u(x, t) = \phi(x + t) + \psi(x - t) \quad (15)$$

remains true even for distributional solutions of the 1D wave equation.

## 2. Spherical means and Euler-Poisson-Darboux equation.

The case  $n \geq 2$  is much more complicated. The idea is to reduce the wave equation to a 1D equation which can be solved explicitly. The reduction is fulfilled through introducing the following auxiliary functions.

Let  $u = u(x, t)$ . We define at each  $x \in \mathbb{R}^n$ ,

$$U(x; r, t) \equiv \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} u(w, t) dS_w, \quad (16)$$

$$G(x; r) \equiv \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} g(w) dS_w, \quad (17)$$

$$H(x; r) \equiv \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} h(w) dS_w. \quad (18)$$

Note that when  $u$  is continuous, we can recover  $u$  from  $U$  by taking  $r \searrow 0$ .

It turns out that  $U(x; r, t)$  as a function of  $r$  and  $t$  satisfies a 1D equation (which can be further transformed into a 1D wave equation in the first quadrant) and enjoys the same regularity as  $u$ .

**Lemma 4. (Euler-Poisson-Darboux equation)** Fix  $x \in \mathbb{R}^n$ . Let  $u(x, t) \in C^m$ ,  $m \geq 2$  solves the wave equation. Then

$$U(x; r, t) \equiv \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} u(w, t) dS_w \quad (19)$$

belongs to  $C^m(\bar{\mathbb{R}}_+ \times [0, \infty))$ , and satisfies

$$U_{tt} - U_{rr} - \frac{n-1}{r} U_r = 0 \quad \mathbb{R}_+ \times (0, \infty); \quad U(r, 0) = G(r), \quad U_t(r, 0) = H(r). \quad (20)$$

**Remark 5.** Notice that  $\partial_{rr} - \frac{n-1}{r} \partial_r$  is just  $\Delta$  in  $\mathbb{R}^n$  with radial symmetry.

**Proof.** Recall that

$$U_r(x; r, t) = \frac{r}{n} \frac{1}{|B_r(x)|} \int_{B_r(x)} \Delta_y u(y, t) dy = \frac{1}{n \alpha(n) r^{n-1}} \int_{B_r(x)} \Delta_y u(y, t) dy \quad (21)$$

This shows  $U \in C^1$ , and we can define  $U_r(x; 0, t) = 0$ .

Differentiating w.r.t  $r$  again,

$$\begin{aligned} U_{rr}(x; r, t) &= \frac{d}{dr} \left[ \frac{1}{n \alpha(n) r^{n-1}} \int_{B_r(x)} \Delta_y u(y, t) dy \right] \\ &= \frac{1-n}{n} \frac{1}{|B_r|} \int_{B_r(x)} \Delta_y u + \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} \Delta_y u. \end{aligned} \quad (22)$$

This shows  $U \in C^2$  and also  $U_{rr}(x; 0, t)$  can be defined.

We further have

$$U_{tt}(x; r, t) = \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} u_{tt} = - \frac{1}{|\partial B_r|} \int_{\partial B_r(x)} \Delta_y u \quad (23)$$

using the equation. □

### 3. $n = 3$ , Kirchhoff's formula.

Let  $U, G, H$  be the spherical means. We set

$$\tilde{U} = rU, \quad \tilde{G} = rG, \quad \tilde{H} = rH. \quad (24)$$

Some calculation yields

$$\tilde{U}_{tt} - \tilde{U}_{rr} = 0 \quad \mathbb{R}_+ \times (0, \infty); \quad \tilde{U} = \tilde{G}, \quad \tilde{U}_t = \tilde{H}. \quad (25)$$

**Remark 6.** Note that here we used the fact that  $n = 3$ .

Thus we need to solve the wave equation in the first quadrant.

**Example 7.** Consider the wave equation in the first quadrant:

$$u_{tt} - u_{xx} = 0, \quad x > 0, t > 0; \quad u(x, 0) = g(x), \quad u_t(x, 0) = h(x), \quad u = 0 \text{ for } x = 0, t > 0. \quad (26)$$

Let

$$\tilde{u}(x, t) = \begin{cases} u(x, t) & x > 0 \\ -u(-x, t) & x < 0 \end{cases} \quad (27)$$

and define similarly  $\tilde{g}, \tilde{h}$ . Then it is clear that  $\tilde{u}$  solves the wave equation with initial values  $\tilde{g}, \tilde{h}$ . Thus we have

$$\tilde{u}(x, t) = \frac{1}{2} [\tilde{g}(x+t) + \tilde{g}(x-t)] + \frac{1}{2} \int_{x-t}^{x+t} \tilde{h}(y) dy. \quad (28)$$

Therefore the solution to the original problem is

$$u(x, t) = \begin{cases} \frac{1}{2} [g(x+t) + g(x-t)] + \frac{1}{2} \int_{x-t}^{x+t} h(y) dy & x \geq t \geq 0 \\ \frac{1}{2} [g(x+t) - g(t-x)] + \frac{1}{2} \int_{t-x}^{t+x} h(y) dy & t \geq x \geq 0 \end{cases} \quad (29)$$

Now for our purpose, we only need the case  $t \geq r$  (remember that finally we will let  $r \searrow 0$  and recover  $u$  from  $U$ ). In this case

$$\tilde{U}(x; r, t) = \frac{1}{2} [\tilde{G}(r+t) - \tilde{G}(t-r)] + \frac{1}{2} \int_{t-r}^{t+r} \tilde{H}(y) dy. \quad (30)$$

We have

$$\begin{aligned} u(x, t) &= \lim_{r \searrow 0} \frac{\tilde{U}(x; r, t)}{r} \\ &= \tilde{G}'(t) + \tilde{H}(t) \\ &= \frac{\partial}{\partial t} \left( \frac{t}{|\partial B_t|} \int_{\partial B_t(x)} g(w) dS_w \right) + \frac{t}{|\partial B_t|} \int_{\partial B_t(x)} h(w) dS_w. \end{aligned} \quad (31)$$

Further computation yields

$$u(x, t) = \frac{1}{|\partial B_t|} \int_{\partial B_t(x)} [t h(w) + g(w) + \nabla g(w) \cdot (w - x)] dS_w \quad (32)$$

which is the Kirchhoff's formula.

### 4. $n=2$ , Method of descent and Poisson's formula.

It is not possible to simplify as we did in the  $n = 3$  case. Instead, we use the so-called "method of descent", which treats the solution  $u(x, t)$  of the 2D wave equation as a solution to the 3D equation. We set

$$\bar{u}(x_1, x_2, x_3, t) \equiv u(x_1, x_2, t). \quad (33)$$

and define  $\bar{g}, \bar{h}$  similarly.

Using the Kirchoff's formula we have

$$\begin{aligned} u(x, t) &= \bar{u}(\bar{x}, t) \\ &= \frac{1}{|\partial B_t(\bar{x})|} \int_{\partial B_t(\bar{x})} t \bar{h}(\bar{w}) + \bar{g}(\bar{w}) + \nabla_{\bar{x}} \bar{g}(\bar{w}) \cdot (\bar{w} - \bar{x}) \, dS_{\bar{w}}. \end{aligned} \quad (34)$$

where  $\bar{x} = (x, x_3)$  and  $B_t(\bar{x})$  is the ball in  $\mathbb{R}^3$ .

From definitions of the various bar-ed functions, we have

$$u(x, t) = \frac{1}{4\pi t^2} \int_{\partial B_t(\bar{x})} t h(y) + g(y) + \nabla_y g(y) \cdot (y - x) \, dS_{\bar{w}} \quad (35)$$

where  $\bar{w} = \left( y, \pm \sqrt{t^2 - |y|^2} \right)$ .

Finally, let  $D_t(x)$  denote the ball in  $\mathbb{R}^2$  centered at  $x$  with radius  $t$ , we have

$$\begin{aligned} u(x, t) &= \frac{2}{4\pi t^2} \int_{D_t(x)} \frac{t h(y) + g(y) + \nabla_y g(y) \cdot (y - x)}{\left(1 - \frac{|y-x|^2}{t^2}\right)^{1/2}} \, dy \\ &= \frac{1}{2} \frac{1}{|D_t|} \int_{D_t(x)} \frac{t g(y) + t^2 h(y) + t \nabla g \cdot (y - x)}{(t^2 - |y-x|^2)^{1/2}} \, dy. \end{aligned} \quad (36)$$

This is the Poisson's formula.

**Remark 8. (Huygens' Principle)** We notice that the behavior of the solutions for the 2D and 3D wave equations are drastically different. In 2D,  $u(x, t)$  depends on initial data in the whole ball  $D_t(x)$  while in 3D it only depends on the data on the boundary of the ball  $B_t(x)$ . Or equivalently, in 3D the effect of a vibration is only felt at the front of its propagation while in 2D it is felt forever after the front passed.<sup>1</sup> This is the so-called Huygens' principle.

**Remark 9.** For general  $n$ , we define

$$\tilde{U}(r, t) = \left( \frac{1}{r} \partial_r \right)^{k-1} (r^{2k-1} U(x; r, t)) \quad (37)$$

and define  $\tilde{G}, \tilde{H}$  accordingly. Some calculation yields the solution

$$u(x, t) = \frac{1}{\gamma_n} \left[ \left( \frac{\partial}{\partial t} \right) \left( \frac{1}{t} \frac{\partial}{\partial t} \right)^{\frac{n-3}{2}} \left( t^{n-2} \frac{1}{|\partial B_t|} \int_{\partial B_t(x)} g \, dS \right) + \left( \frac{1}{t} \frac{\partial}{\partial t} \right)^{\frac{n-3}{2}} \left( t^{n-2} \frac{1}{|\partial B_t|} \int_{\partial B_t(x)} h \, dS \right) \right] \quad (38)$$

for  $n$  odd, where  $\gamma_n = 1 \cdot 3 \cdot \dots \cdot (n-2)$ . Then the method of descent yields

$$u(x, t) = \frac{1}{\gamma_n} \left[ \left( \frac{\partial}{\partial t} \right) \left( \frac{1}{t} \frac{\partial}{\partial t} \right)^{\frac{n-2}{2}} \left( \frac{t^n}{|B_t|} \int_{B_t} \frac{g(y) \, dy}{(t^2 - |y-x|^2)^{1/2}} \right) + \left( \frac{1}{t} \frac{\partial}{\partial t} \right)^{\frac{n-2}{2}} \left( \frac{t^n}{|B_t|} \int_{B_t} \frac{h(y) \, dy}{(t^2 - |y-x|^2)^{1/2}} \right) \right].$$

for  $n$  even, where  $\gamma_n = 2 \cdot 4 \cdot \dots \cdot (n-2) \cdot n$ .

See L. C. Evans **Partial Differential Equations** pp. 75–80 for details.

**Remark 10. (Nonhomogeneous problem)** For the nonhomogeneous problem

$$\square u = f, \quad u = 0, \quad u_t = 0, \quad (39)$$

we use the Duhamel's principle, obtaining

$$u(x, t) = \int_0^t u(x, t; s) \, ds \quad (40)$$

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1. If  $u(x, t)$  also depends on data in the whole ball  $B_t(x)$  in 3D, we would not be able to clearly hear anything!

where  $u(x, t; s)$  solves

$$u_{tt} - \Delta u = 0, \quad u(x, s; s) = 0, \quad u_t(x, s; s) = f(\cdot, s). \quad (41)$$

In particular, we have

–  $n = 1$ :

$$u(x, t) = \frac{1}{2} \int_0^t \int_{x-s}^{x+s} f(y, t-s) \, dy \, ds. \quad (42)$$

–  $n = 3$ :

$$u(x, t) = \frac{1}{4\pi} \int_{B_t(x)} \frac{f(y, t-|y-x|)}{|y-x|} \, dy. \quad (43)$$

Here the integrand is called the “retarded potential”.

### Exercises.

**Exercise 1.** Prove the following theorem for 1D wave equations:

**Theorem.** Assume  $g \in C^2(\mathbb{R})$ ,  $h \in C^1(\mathbb{R})$ , define  $u$  by

$$u(x, t) = \frac{1}{2} [g(x+t) + g(x-t)] + \frac{1}{2} \int_{x-t}^{x+t} h(y) \, dy. \quad (44)$$

Then

i.  $u \in C^2(\mathbb{R} \times [0, \infty))$ ;

ii.  $u_{tt} - u_{xx} = 0$  in  $\mathbb{R} \times (0, \infty)$ ;

iii.  $u$  takes the correct boundary values:

$$\lim_{\substack{(x,t) \rightarrow (x_0,0) \\ t > 0}} u(x, t) = g(x_0); \quad (45)$$

$$\lim_{\substack{(x,t) \rightarrow (x_0,0) \\ t > 0}} u_t(x, t) = h(x_0). \quad (46)$$

**Exercise 2. (Equipartition of energy)** Let  $u \in C^2(\mathbb{R} \times [0, \infty))$  solve the initial-value problem for the 1D wave equation

$$u_{tt} - u_{xx} = 0, \quad u = g, \quad u_t = h; \quad (47)$$

Suppose  $g, h$  have compact support. Let

$$k(t) \equiv \frac{1}{2} \int_{-\infty}^{\infty} u_t^2(x, t) \, dx, \quad p(t) \equiv \frac{1}{2} \int_{-\infty}^{\infty} u_x^2(x, t) \, dx \quad (48)$$

be the kinetic and potential energy. Prove that for  $t$  large enough,  $k(t) = p(t) = \text{constant}$ .

(Hint: Use d’Alembert’s formula to compute the energy directly.)